

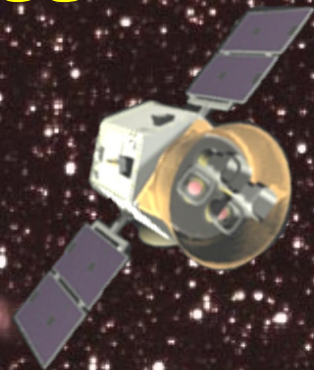
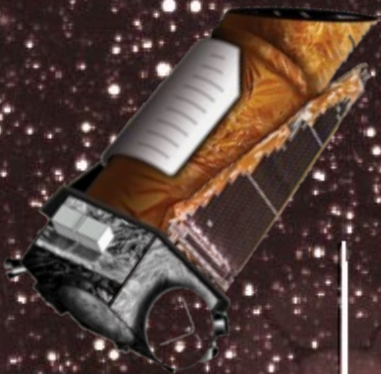


# ***Expanding the Search for Earth-size Planets Orbiting Sun-like Stars by Compressing Image Data Onboard Space Telescopes***

**Jon M. Jenkins  
NASA Ames Research Center**

**Wednesday August 9, 2023**

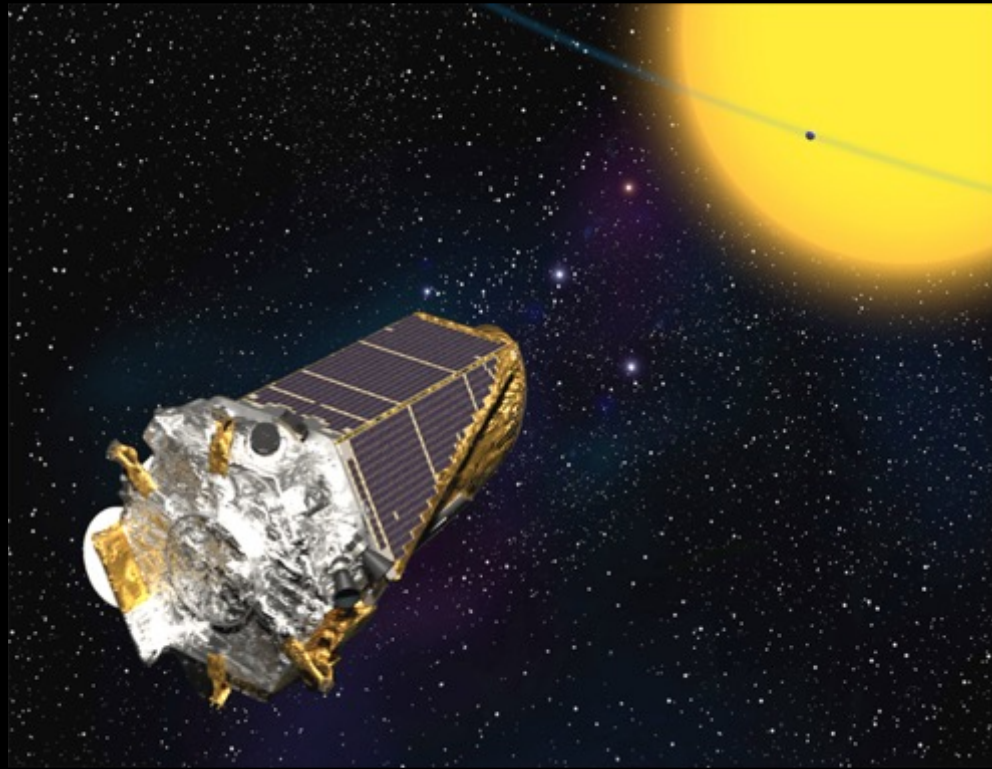
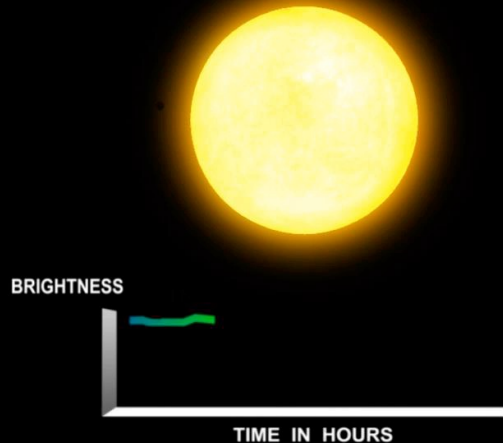
**Flash Memory Summit 2023  
Santa Clara Convention Center**



# The *Kepler* Mission

How many stars like the Sun have  
Earth-like planets orbiting them?

Kepler searches for transiting  
planets



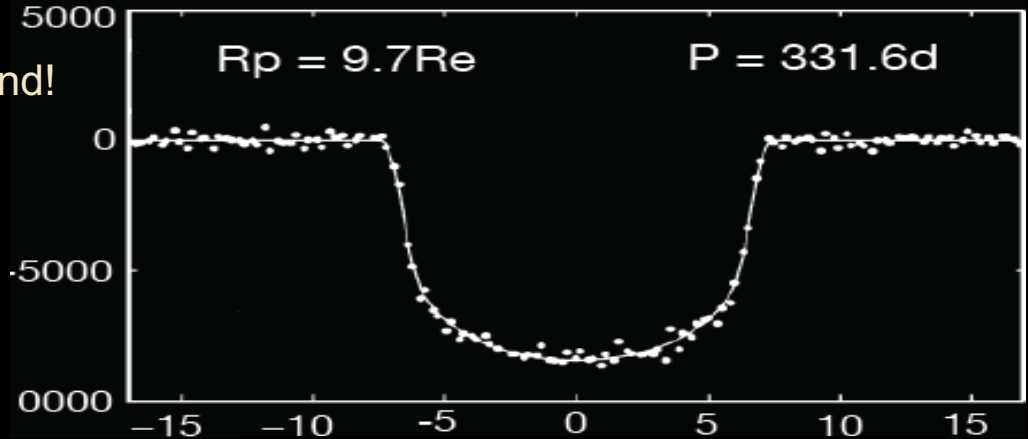


# How Hard is it to Find Good Planets?

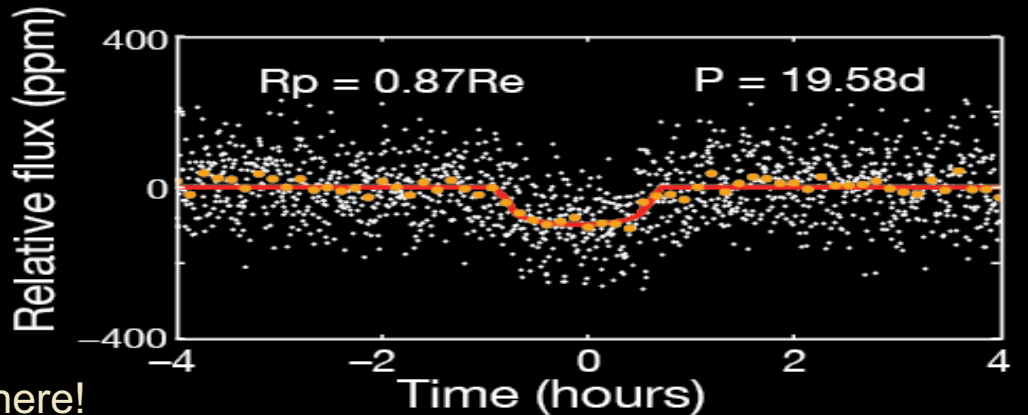


Easy to find from the ground!

Jupiter (~1%)

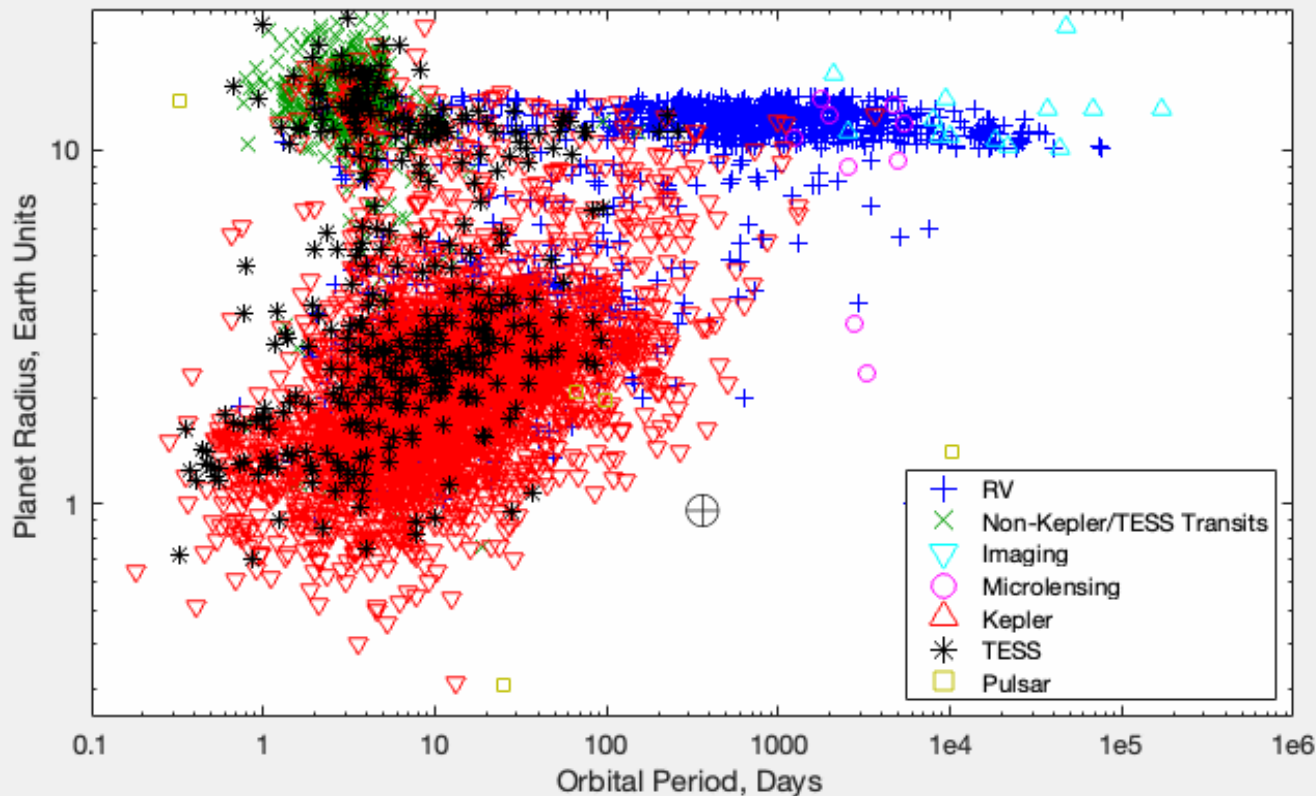


Earth (~0.01%)



Hard to find from anywhere!

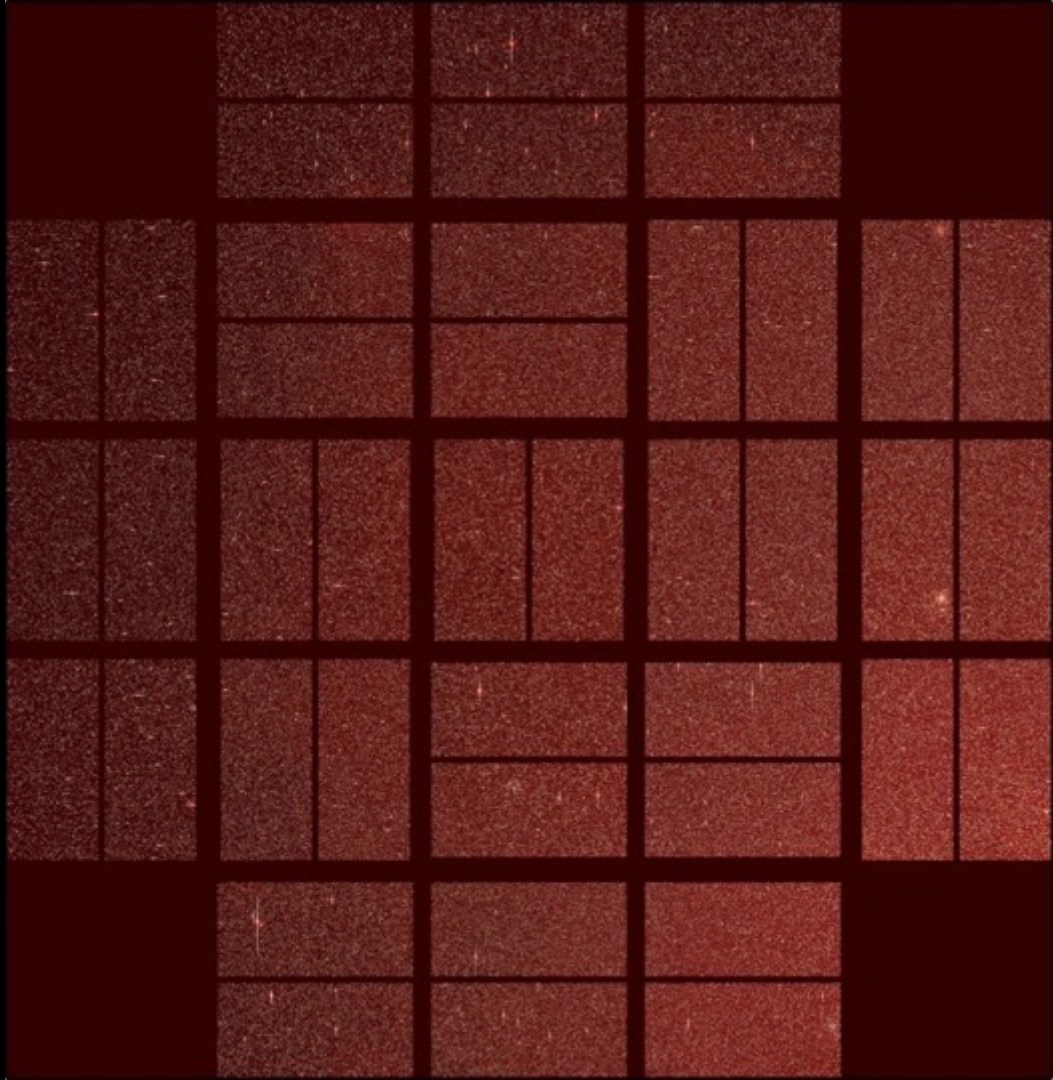
# Exoplanet Discoveries – Over 5470 and Counting!



75% of exoplanets  
have been  
discovered using  
the transit method!

# First Light Image

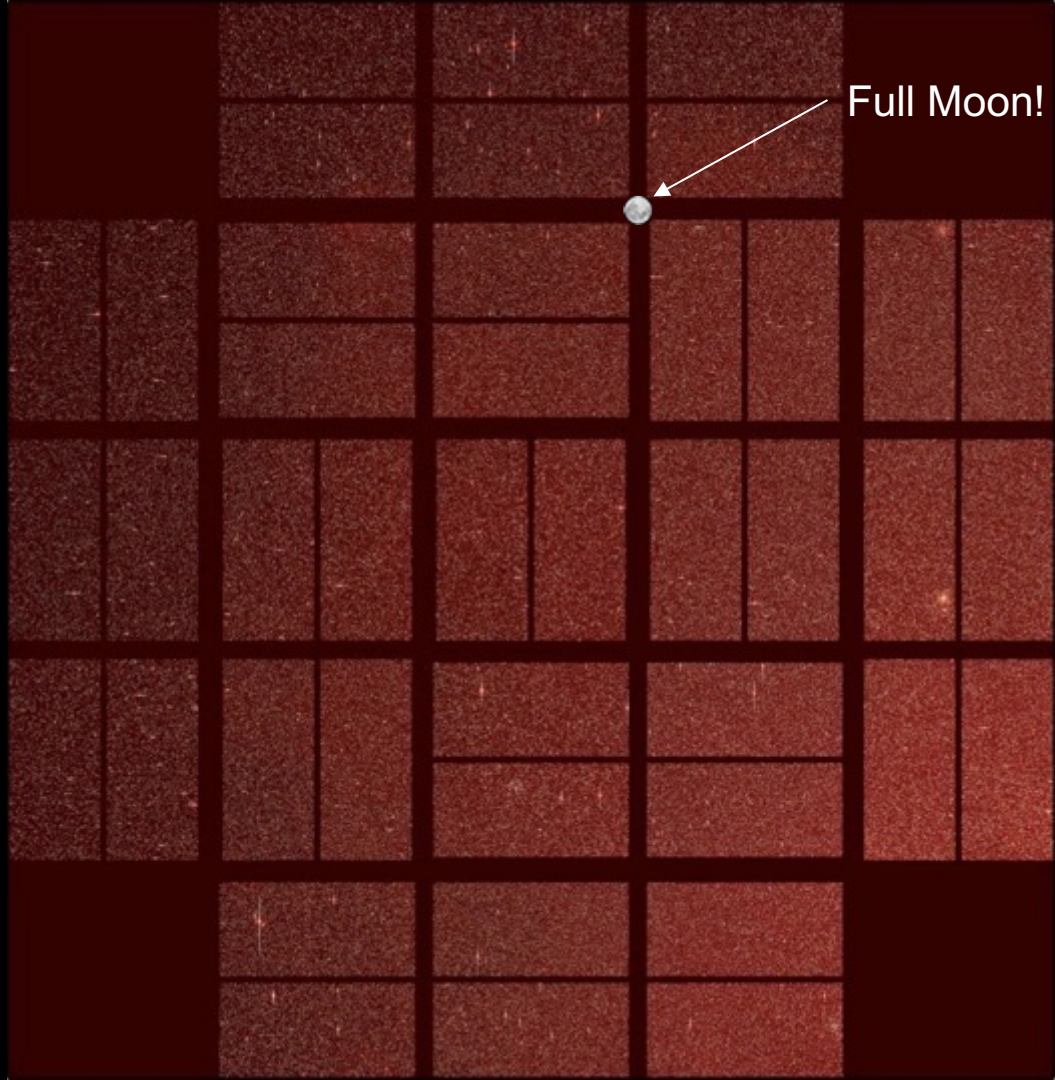
Launched  
March 7 2009





# First Light Image

Launched  
March 7 2009



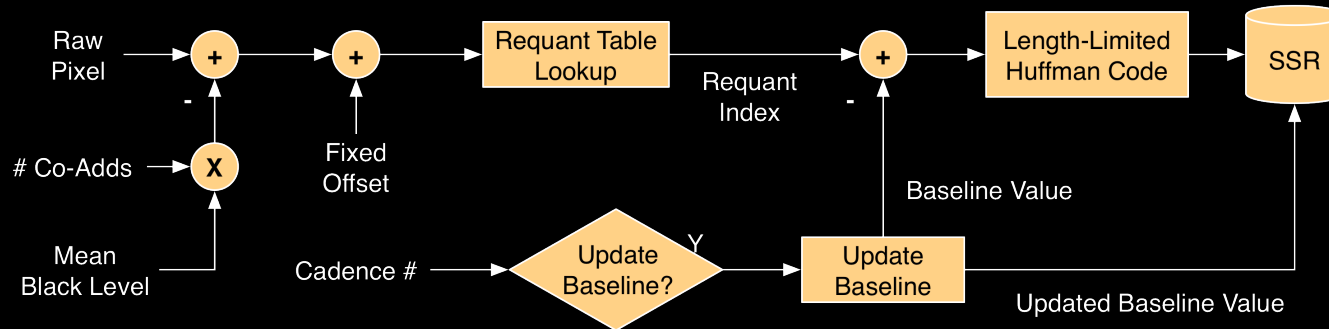
Full Moon!

## Key Requirements:

- Collect and store image data for up to 170,000 stars for up to 66+ days
- Downlink 31 days of data in less than 24 hours

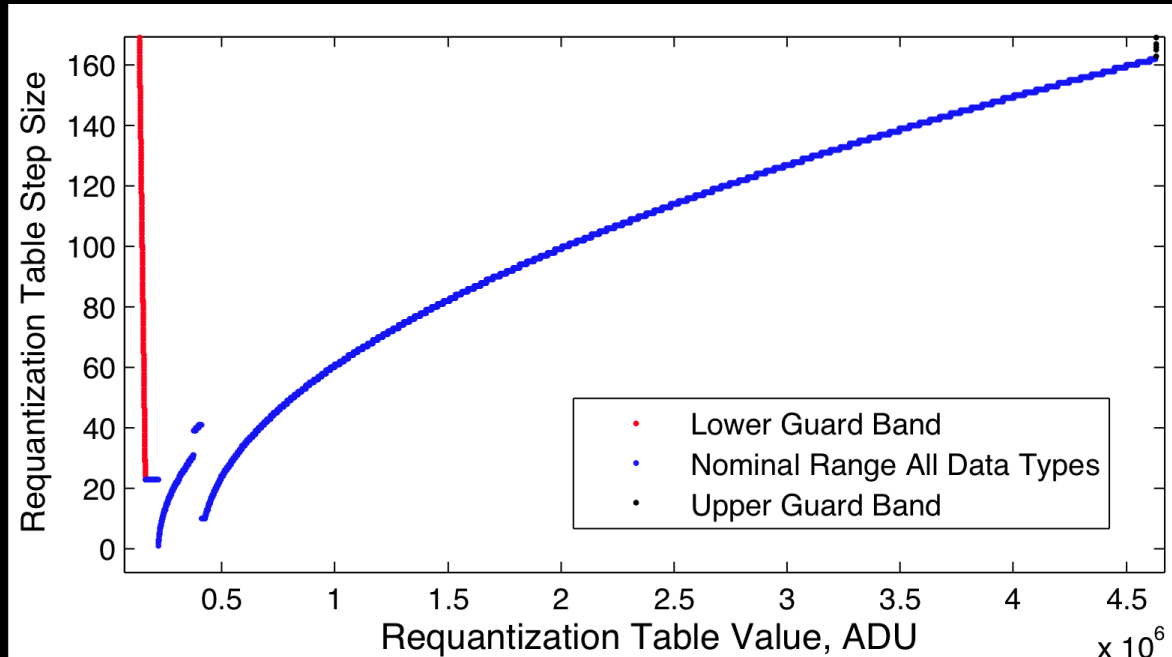
## Solution:

- Collect **only** pixels of interest
- Compand data to manage quantization noise as a term in noise budget
- Detrend the data to centralize the measurements to be transmitted
- Entropically encode the data before CCSDS packetization (must be robust)



# Step 1: Requantization

- Manage quantization noise in noise budget –  $\frac{1}{4}$  of intrinsic measurement noise
- Must account for dispersion of read noise, gain, and bias across 84 channels
- Raw pixel measurements are represented with 23 bits
- Requantization compresses the data to 16 bits from 23





## Step 2: Huffman Coding

- Subtract a baseline measurement (changes on a daily basis)
- Entropically encode residuals with a Length-Limited Huffman Code
- Compresses the data to 4.4 bits/symbol for Kepler
- First difference baseline for TESS compresses data to  $\sim 3.5$  bits/symbol

